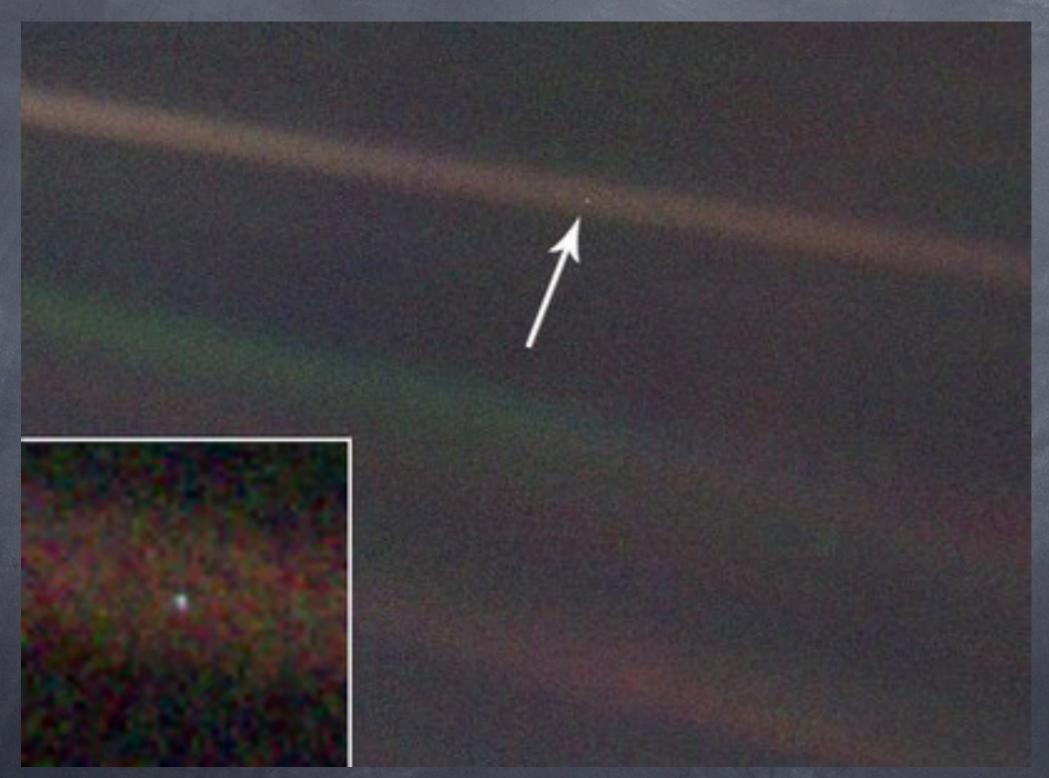
Colours from Earth an inward perspective

Bob Fosbury (The Bob)
@
Villa Aureli, May 2011

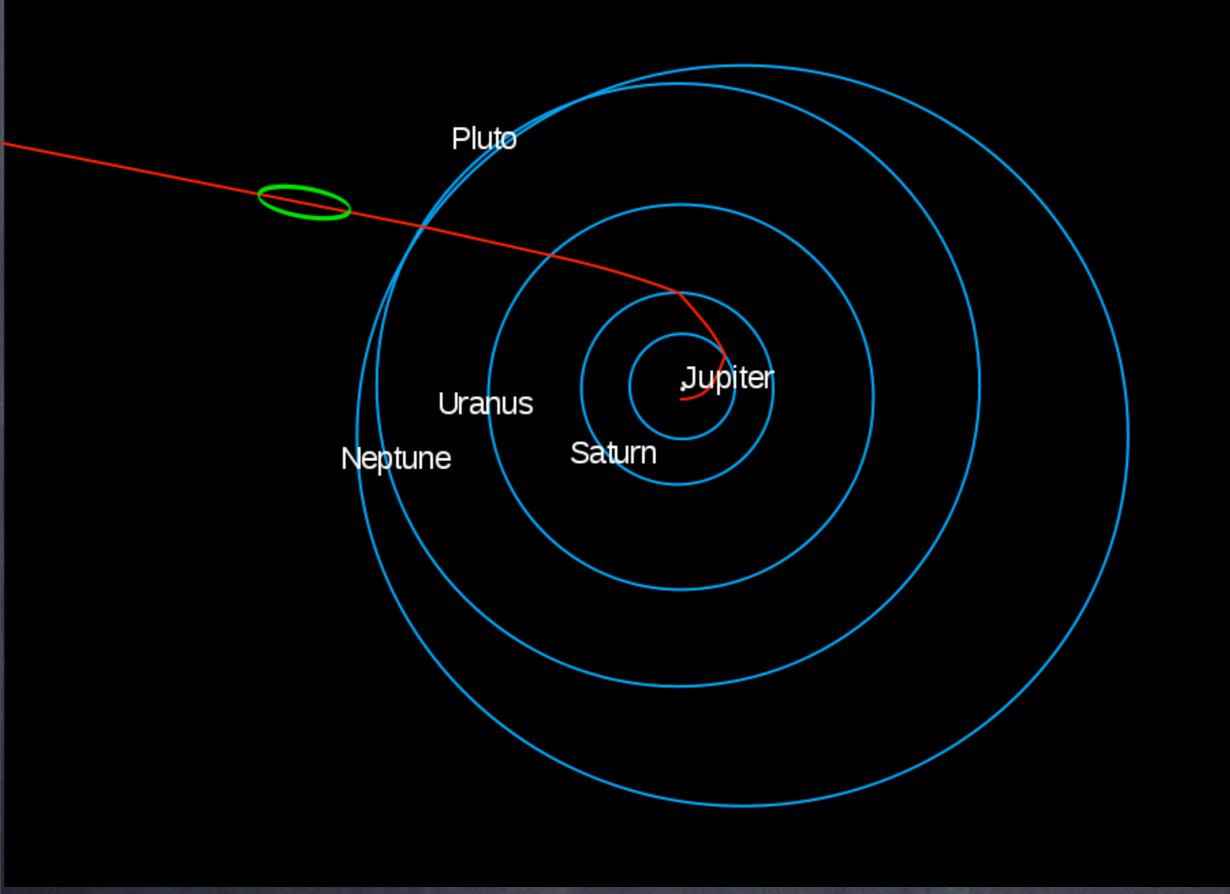


Bob Fosbury (The Bob)
@
Villa Aureli, May 2011

The "Pale Blue Dot"



...taken in 1990 by Voyager 1 from 6.1 billion km



...taken in 1990 by Voyager 1 from 6.1 billion km

From behind Saturn



Cassini Wide Angle Camera, 15 Sep. 2006

From behind Saturn



Cassini Wide Angle Camera, 15 Sep. 2006

From behind Saturn



Cassini Wide Angle Camera, 15 Sep. 2006

From a neighbour

Mars Global Surveyor Mars Orbiter Camera on 8 May 2003 from 139 million km



From a neighbour



Earthrise



... from the Moon: Apollo 8, December 1968

From Earth orbit



Animation by Martin Kornmesser

The "Blue Planet"
The "Pale Blue Dot"

Why blue?

The simple answer: because of air and water But we can look a bit deeper...



The simple answer: because of air and water But we can look a bit deeper...



The simple answer: because of air and water But we can look a bit deeper...

Water is blue

Unlike many blues in nature, water **is** actually blue

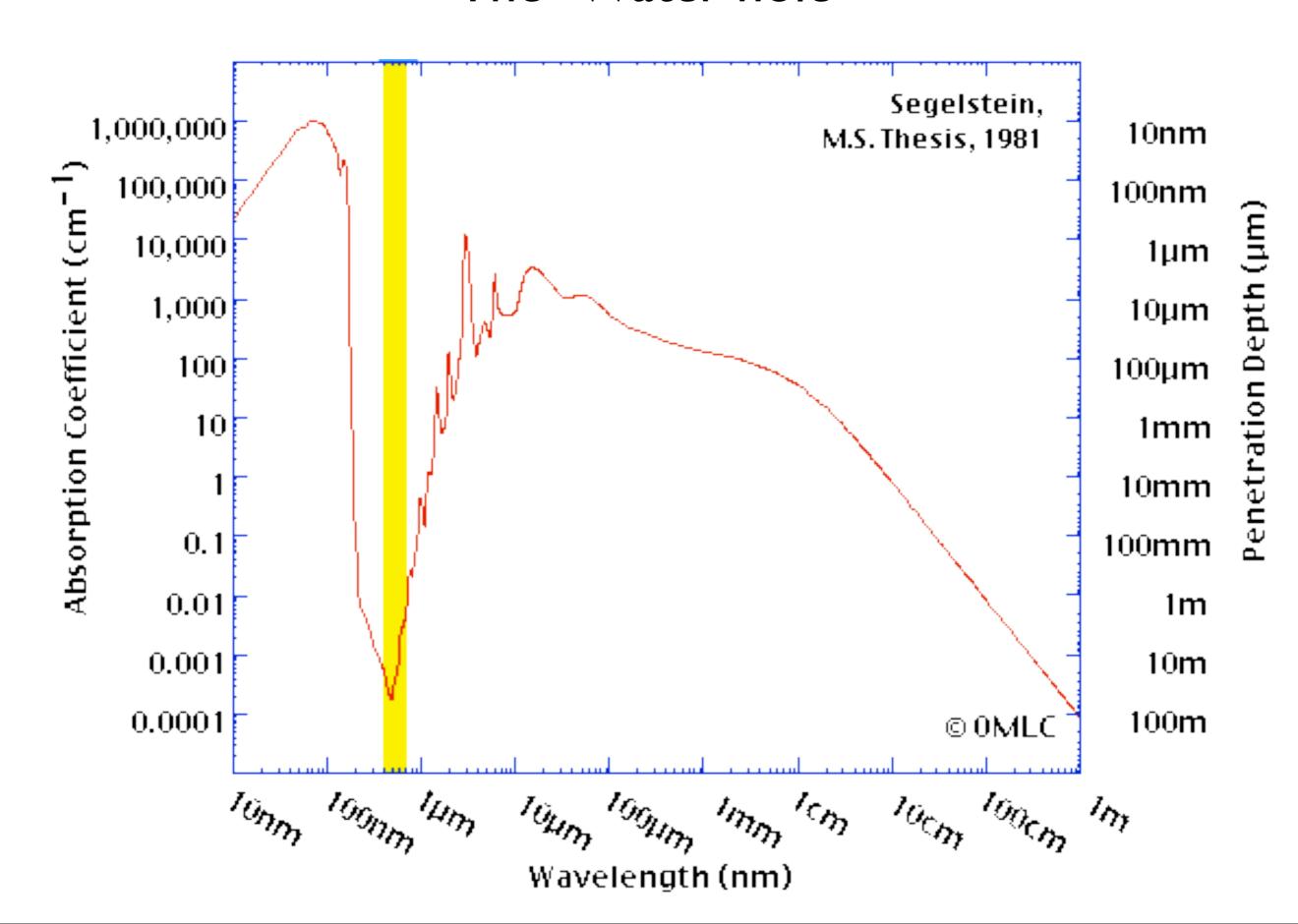
Water molecules absorb red light; in fact water absorbs pretty much all electromagnetic radiation except blue light and metre-length radio waves





Iceberg, Glacier Bay, Alaska

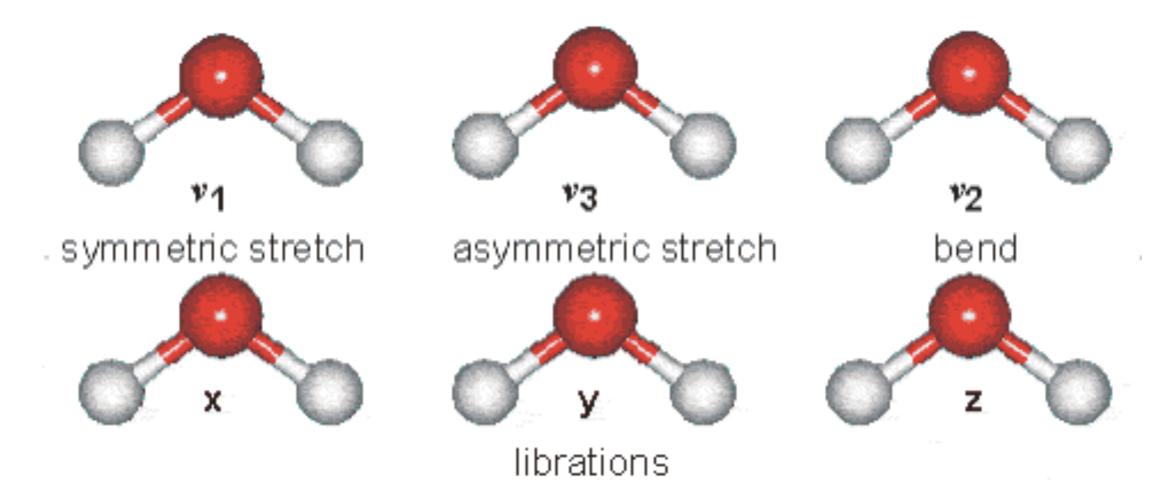
The "Water hole"



Assignment of the IR vibrational absorption spectrum of liquid water

Wavelength	cm ⁻¹	Assignment	Wavelength	cm ⁻¹	Assignment	
0.2 mm	50	intermolecular bend	1470 nm	6800	$av_1 + bv_3$; $a+b=2$	
55 μm	183.4	intermolecular stretch	1200 nm	8330	$av_1 + v_2 + bv_3$; $a+b=2$	
25 μm	395.5	L ₁ , librations	970 nm	10310	$av_1 + bv_3$; $a+b=3$	
15 μm	686.3	L ₂ , librations	836 nm	11960	$av_1 + v_2 + bv_3$; $a+b=3$	
6.08 µm	1645	v ₂ , bend	739 nm	13530	$av_1 + bv_3$; $a+b=4$	
4.65 μm	2150	$v_2 + L_2 \frac{b}{}$	660 nm	15150	$av_1 + v_2 + bv_3$; $a+b=4$	
3.05 µm	3277	v ₁ , symmetric stretch	606 nm	16500	$av_1 + bv_3$; $a+b=5$	
2.87 μm	3490	v ₃ , asymmetric stretch	514 nm	19460	$av_1 + bv_3$; $a+b=6$	
1900 nm	5260	$av_1 + v_2 + bv_3$; $a+b=1$	Note that a and b are integers, \geq 0 ms.			

Thanks to martin.chaplin@btinternet.com

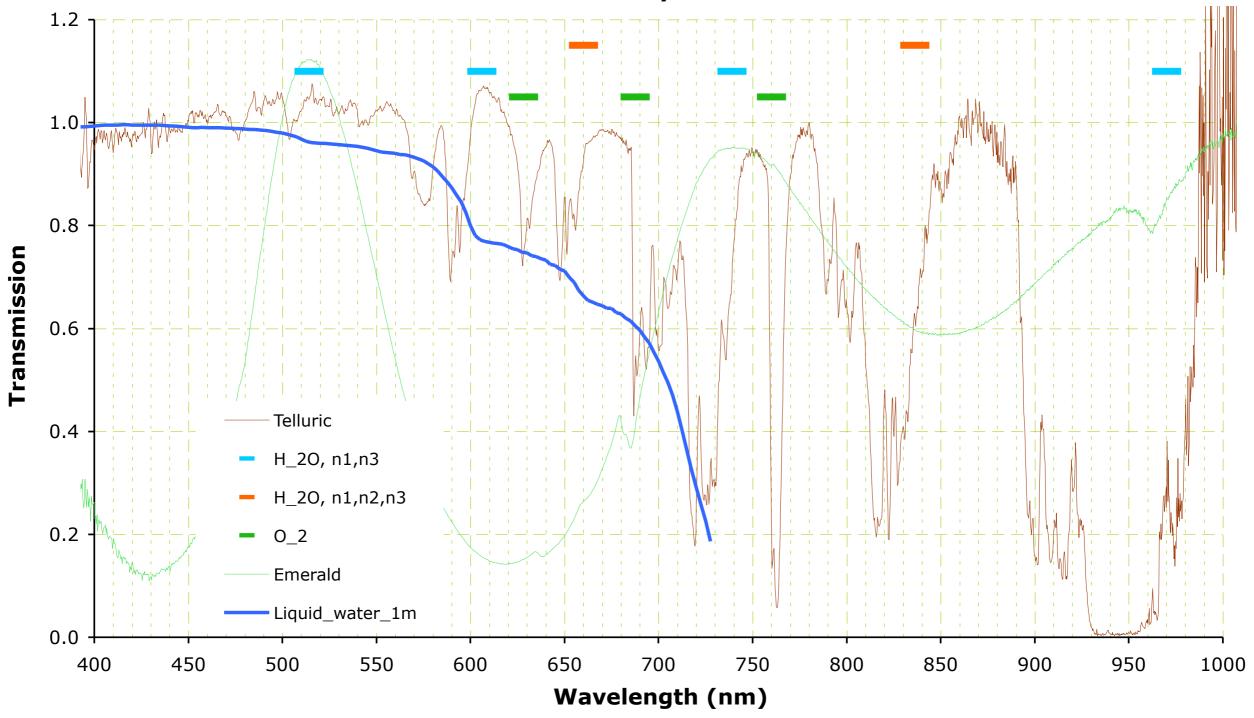


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Thanks to martin.chaplin@btinternet.com

Water absorption in the atmosphere at sunset, in 1m of liquid and in an emerald crystal







The blue colour of heavily overcast skies

Lee et al. 2005

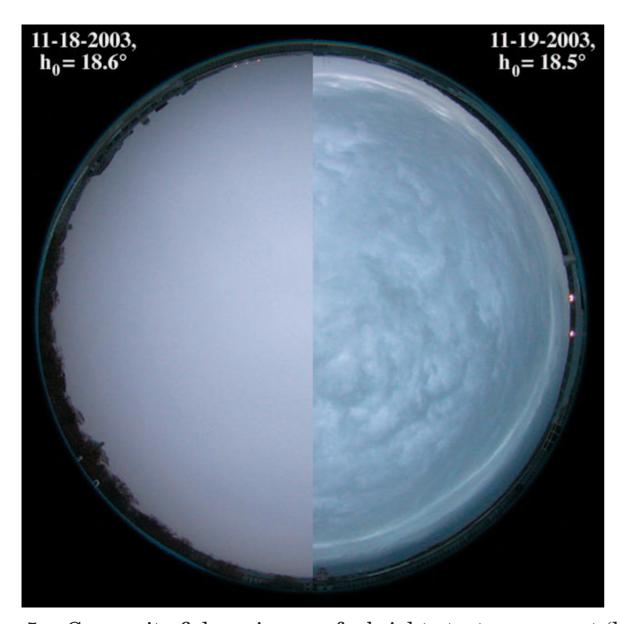


Fig. 5. Composite fisheye image of a bright stratus overcast (left half, $E=24.3~\rm W~m^{-2}$) and a much darker stratocumulus overcast (right half, $E\sim3.6~\rm W~m^{-2}$) at USNA on 11-18-03 and 11-19-03, respectively. Note that the darker overcast is distinctly bluer. Although exposures differ, in both original photographs the digital camera's white-balance setting was the same and $h_0\sim18.5^\circ$.

Mostriamo in figura 4.1 il grafico della distribuzione del numero di interazioni

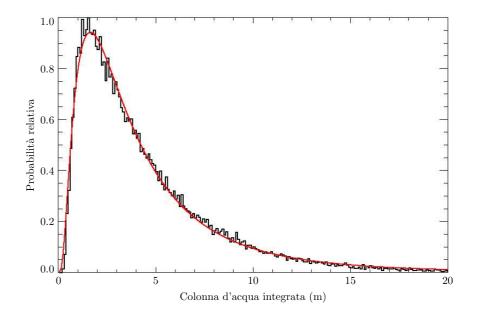


Figura 4.1: Istogramma dei valori della colonna d'acqua. In rosso è rappresentata la curva di fit con andamento log-normale.

Il numero di conteggi inizialmente presente sull'ordinata è stato sostituito con una distribuzione di probabilità normalizzata sul valore di conteggi maggiore (posto quindi uguale a 1). Riportiamo i valori della colonna d'acqua ottenuti dalla simulazione:

$$W$$
 (valor medio) = 4.38 m ,
$$W \mbox{ (mediana)} = 3.32 \mbox{ m} \ ,$$

$$\sigma = 2.77 \mbox{ m} \ . \mbox{ (4.1)}$$

il fit migliore dell'istogramma è dato dalla curva log-normale; la distribuzione log-normale con cui abbiamo fittato l'istogramma presenta una sigma molto alta, cosa che del resto si riscontra nella σ dei nostri dati (eq. 4.1).

Plot for Icm precipitable liquid water

Air is blue

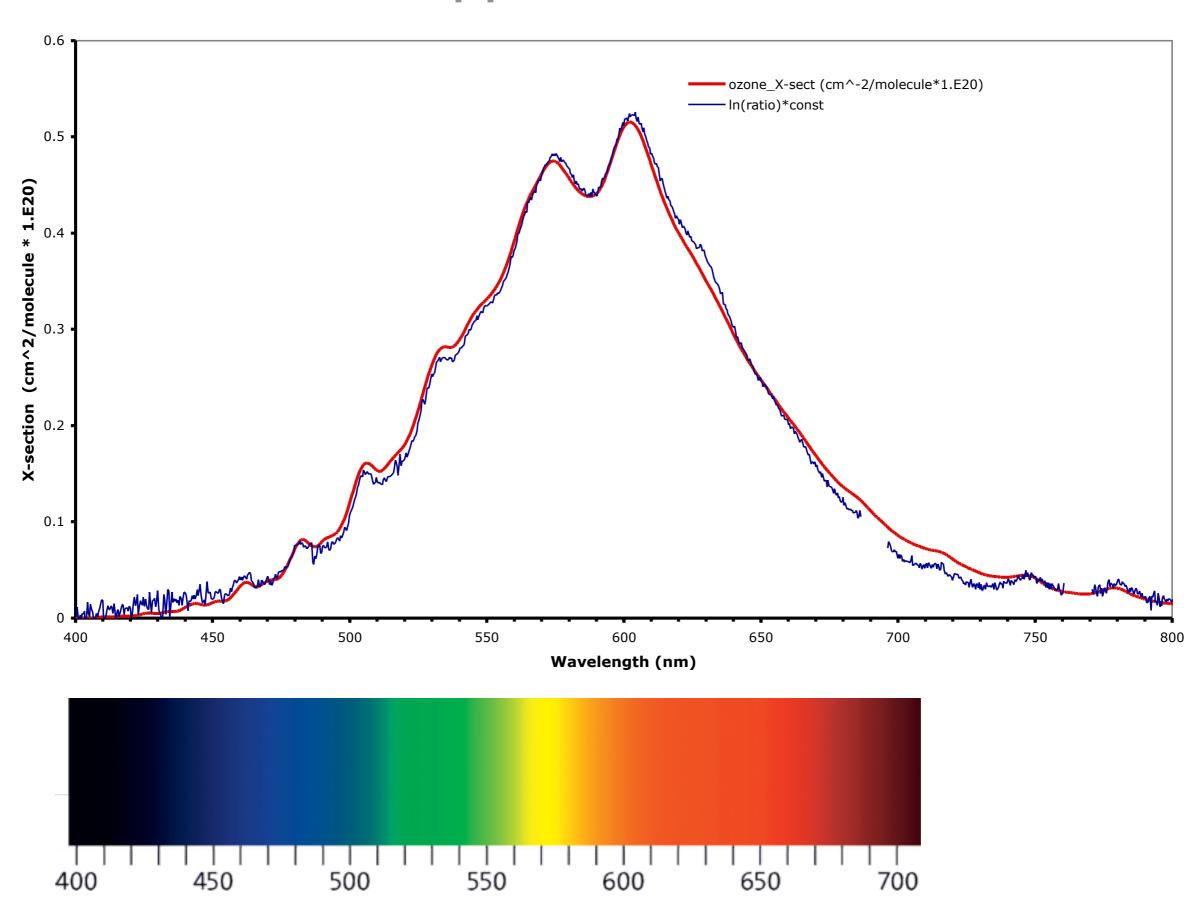
Yes, we all know the clear sky is blue because of Rayleigh scattering from molecules and density fluctuations, but air **is** actually blue as well!

This is because of ozone, the unstable allotrope of oxygen, which exists mostly between 15 and 40 km altitude in the atmosphere



Rayleigh scattering blue Mie scattering white

The Chappuis band of ozone







Lunar eclipse, 16 August 2008 10.00 Sky model Eclipsed Moon 1.00 Model, ozone=0 Ratios 0.10

0.01

350

400

450

500

550

600

650

Wavelength (nm)

700

750

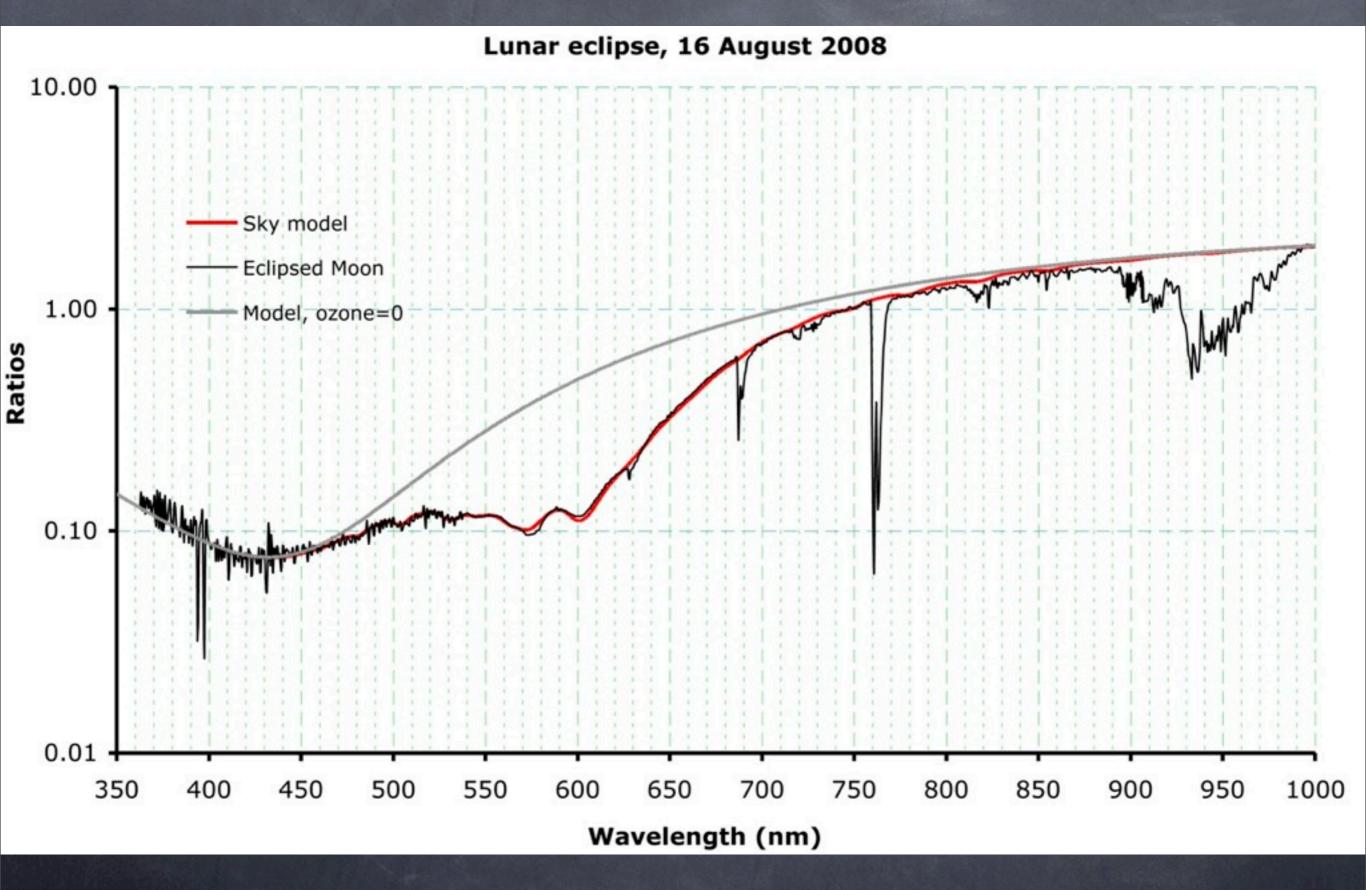
800

850

900

950

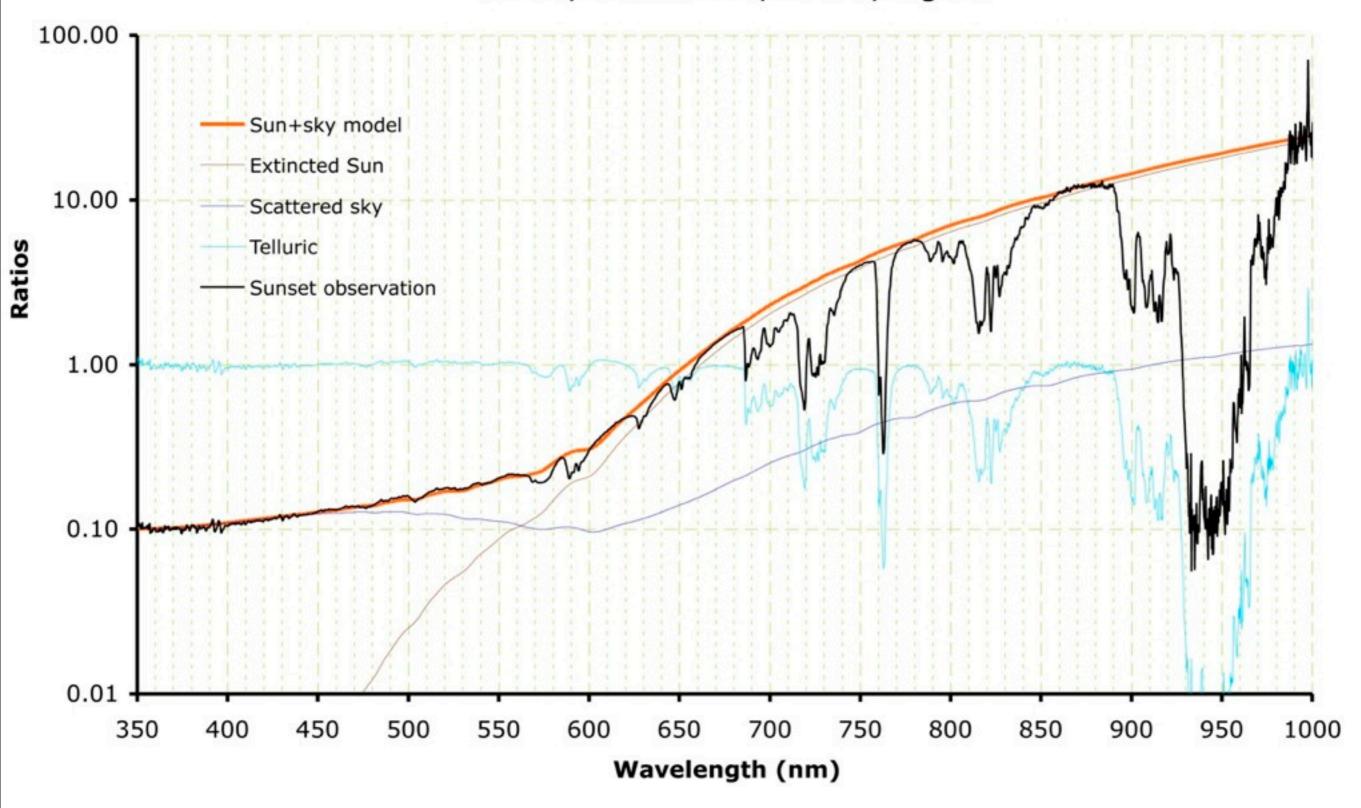
1000





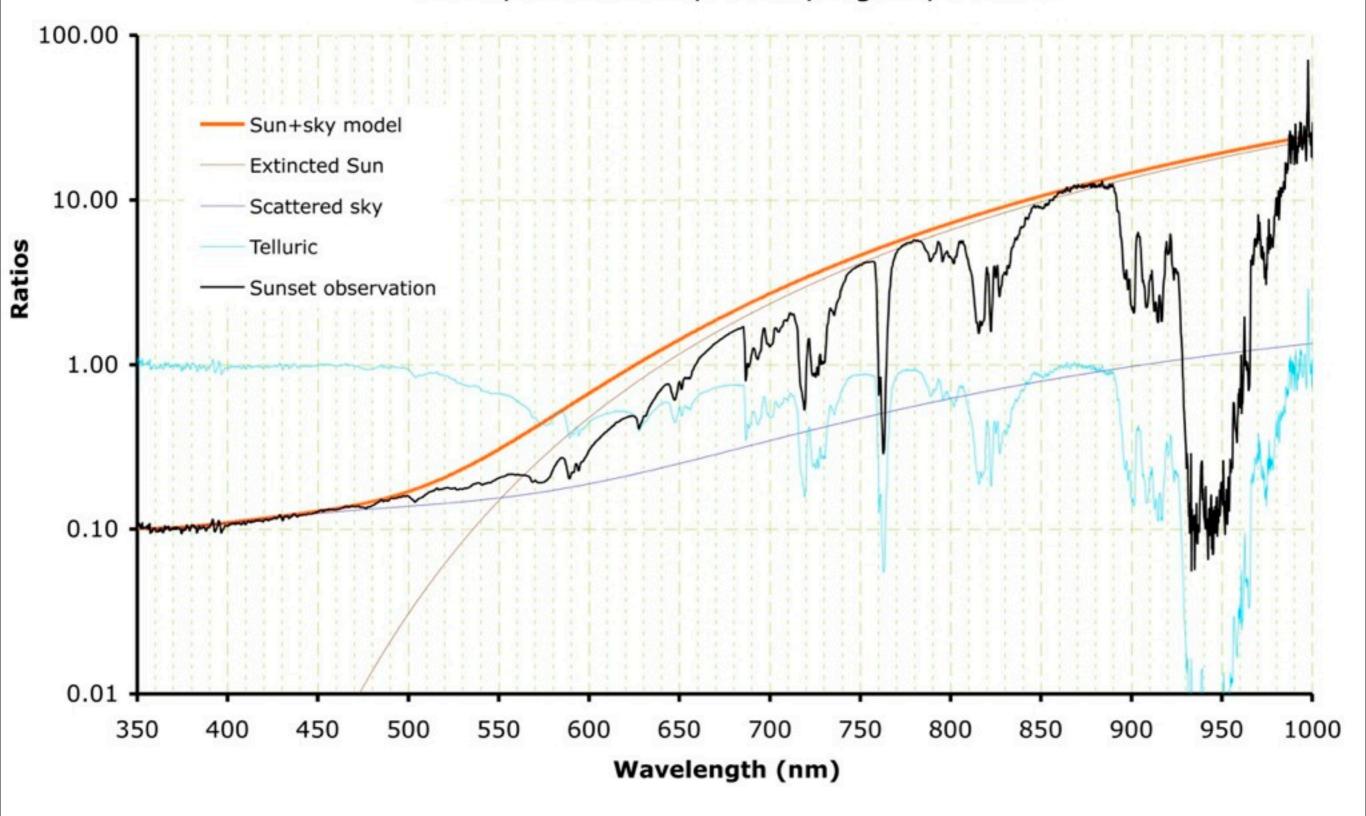
Sunset spectra

Sunset, 27 June 2010, Norfolk, England



Sunset spectra

Sunset, 27 June 2010, Norfolk, England; ozone=0



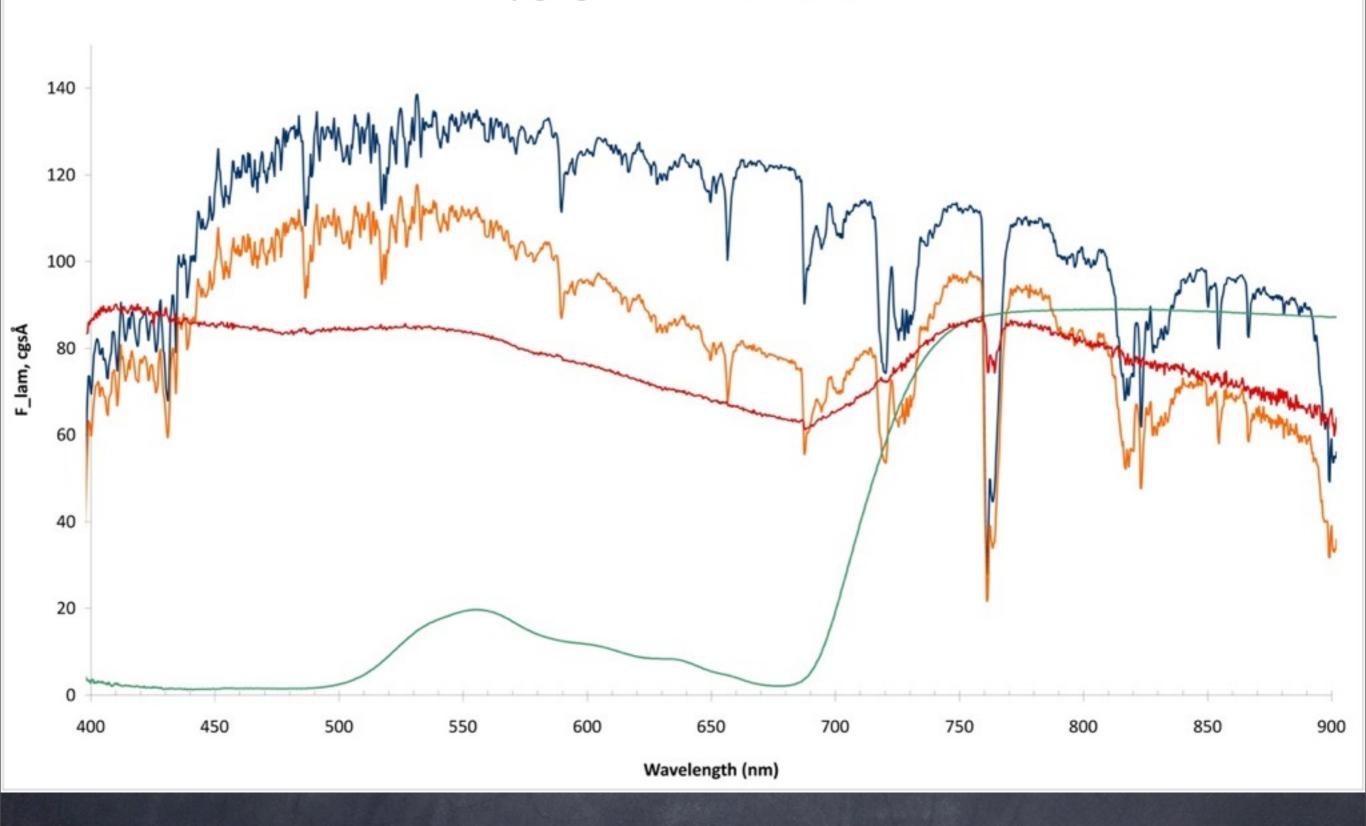
Sunset sky (model)

No ozone

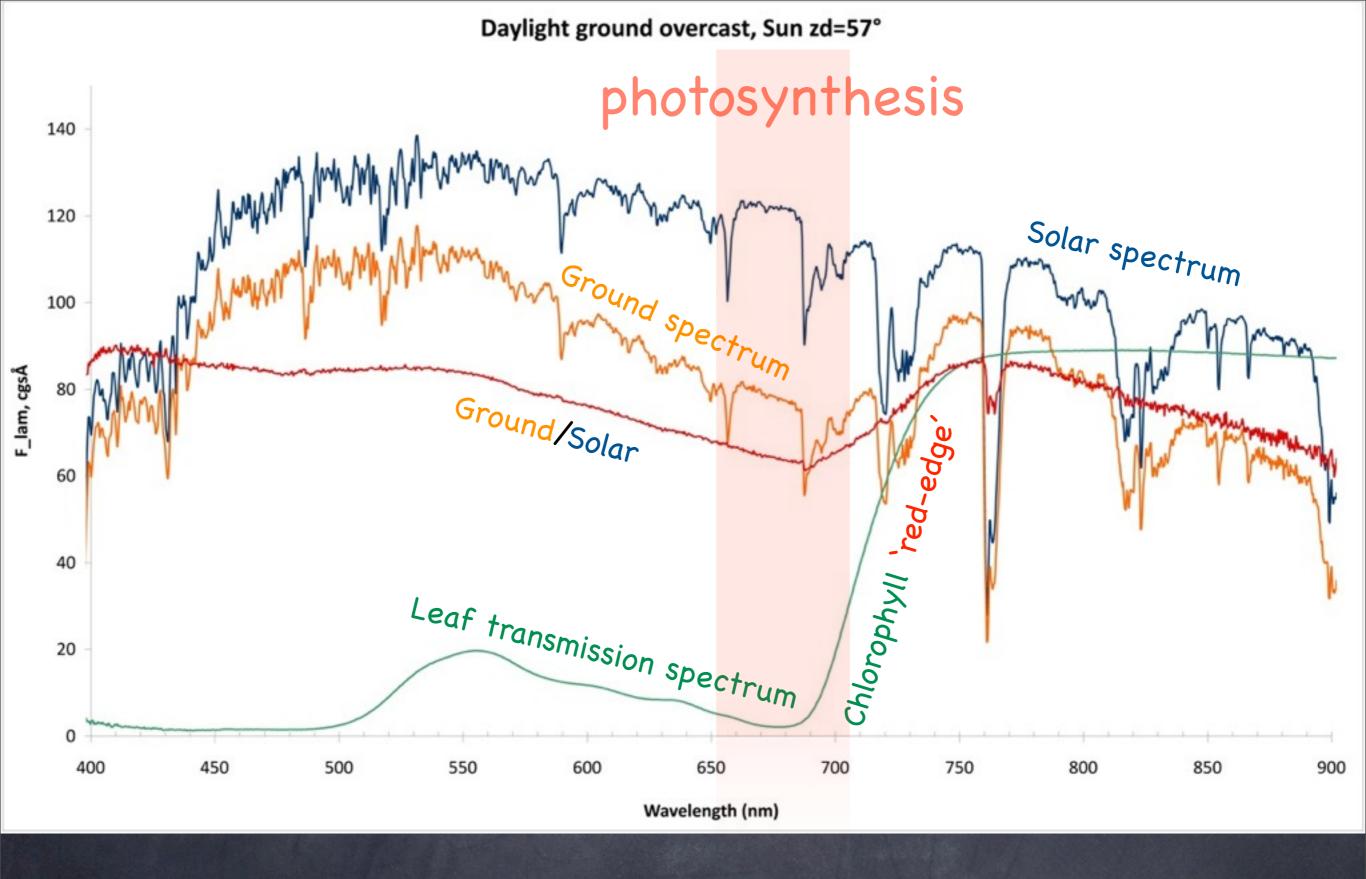
The Green Planet

Chlorophyll in action





Spectrum of the ground: mixed buildings, grass and trees Seeing the chlorophyll 'red-edge'



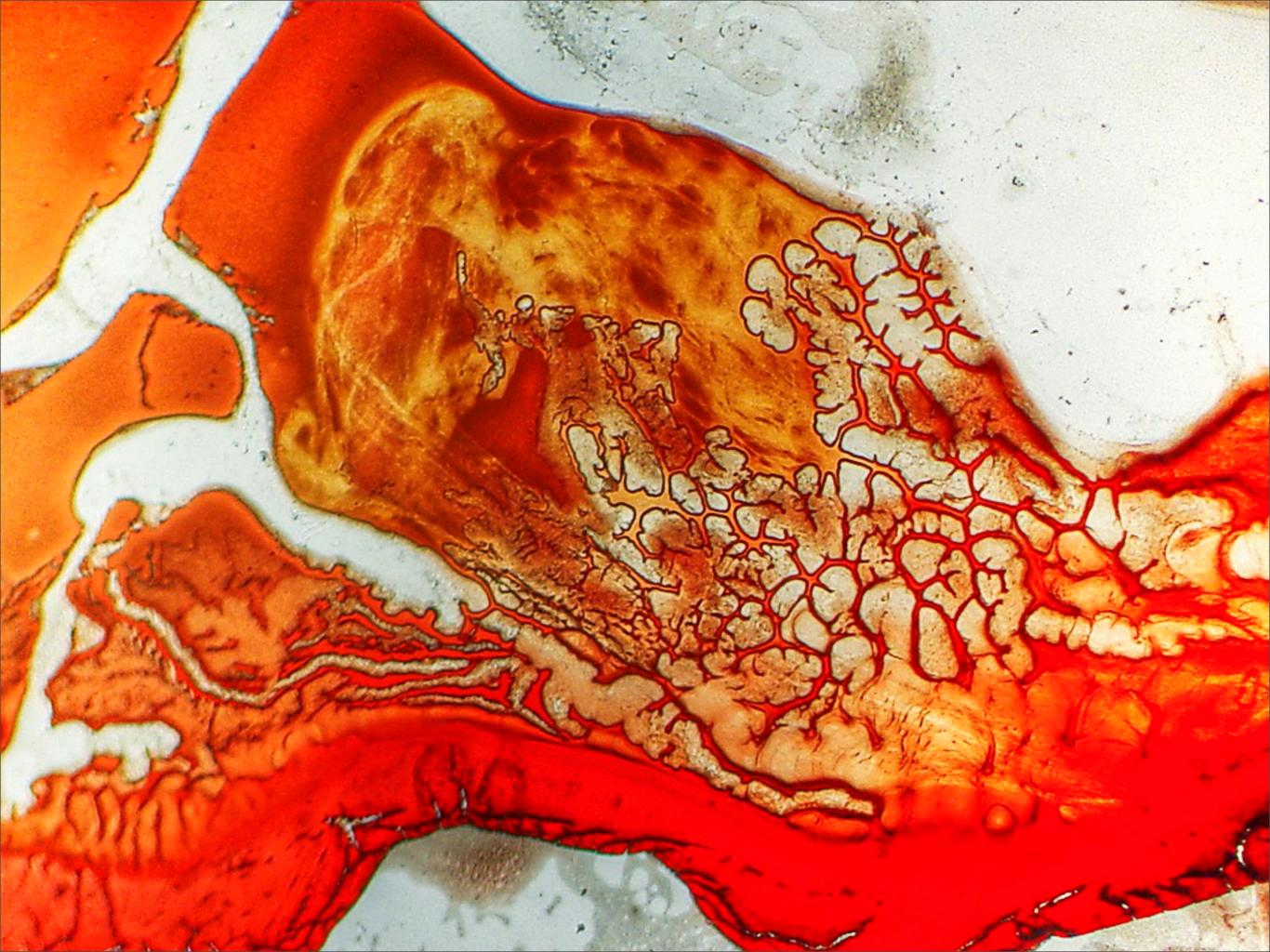
Spectrum of the ground: mixed buildings, grass and trees Seeing the chlorophyll 'red-edge'



Spectrum of the ground: mixed buildings, grass and trees Seeing the chlorophyll 'red-edge'

Chlorophyll, like haemaglobin, is a 'porphyrin' and there are other examples of this class that produce colours we see

The brown of a hen's eggshell is actually a white matrix speckled with purple protoporphyrin and, like chlorophyll, fluoresces red under UV light



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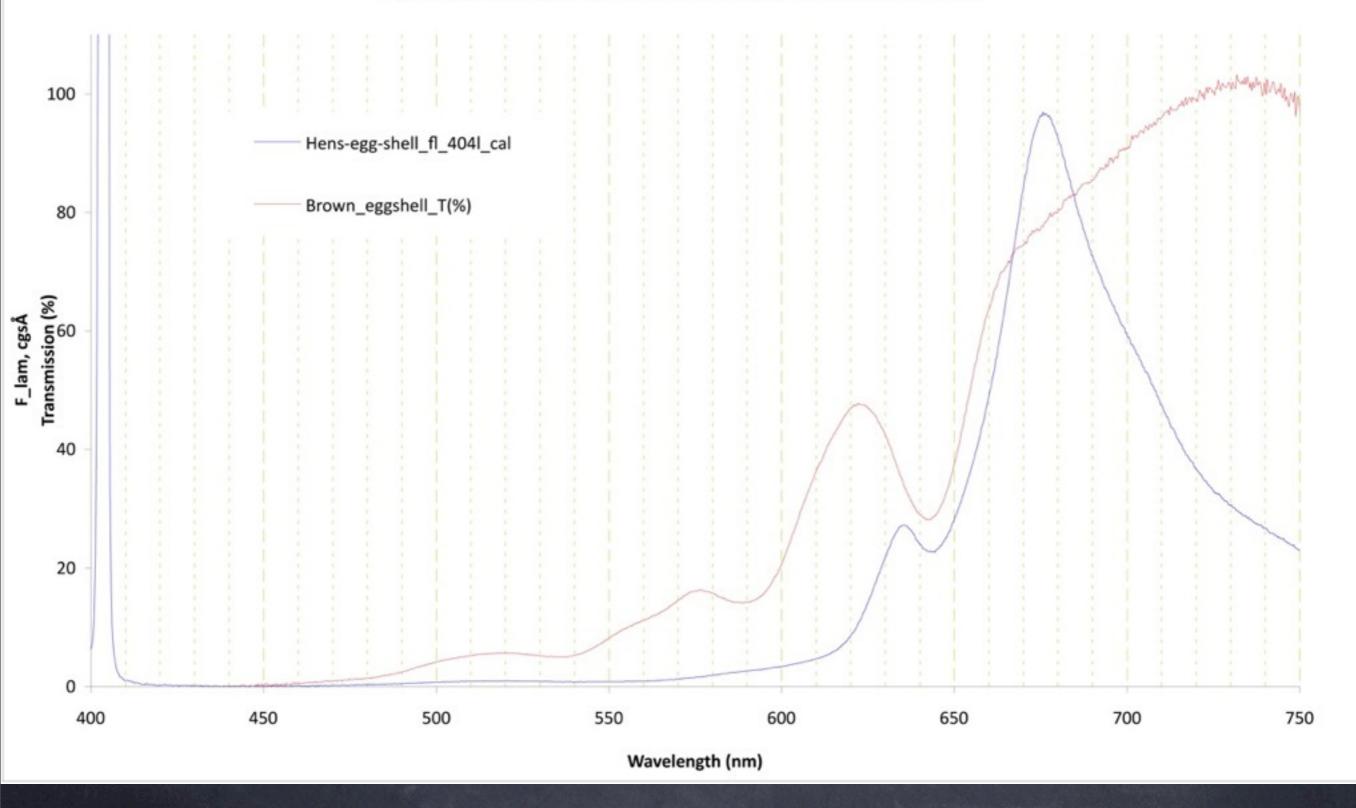
luoresces

that stly blue!) nisms in

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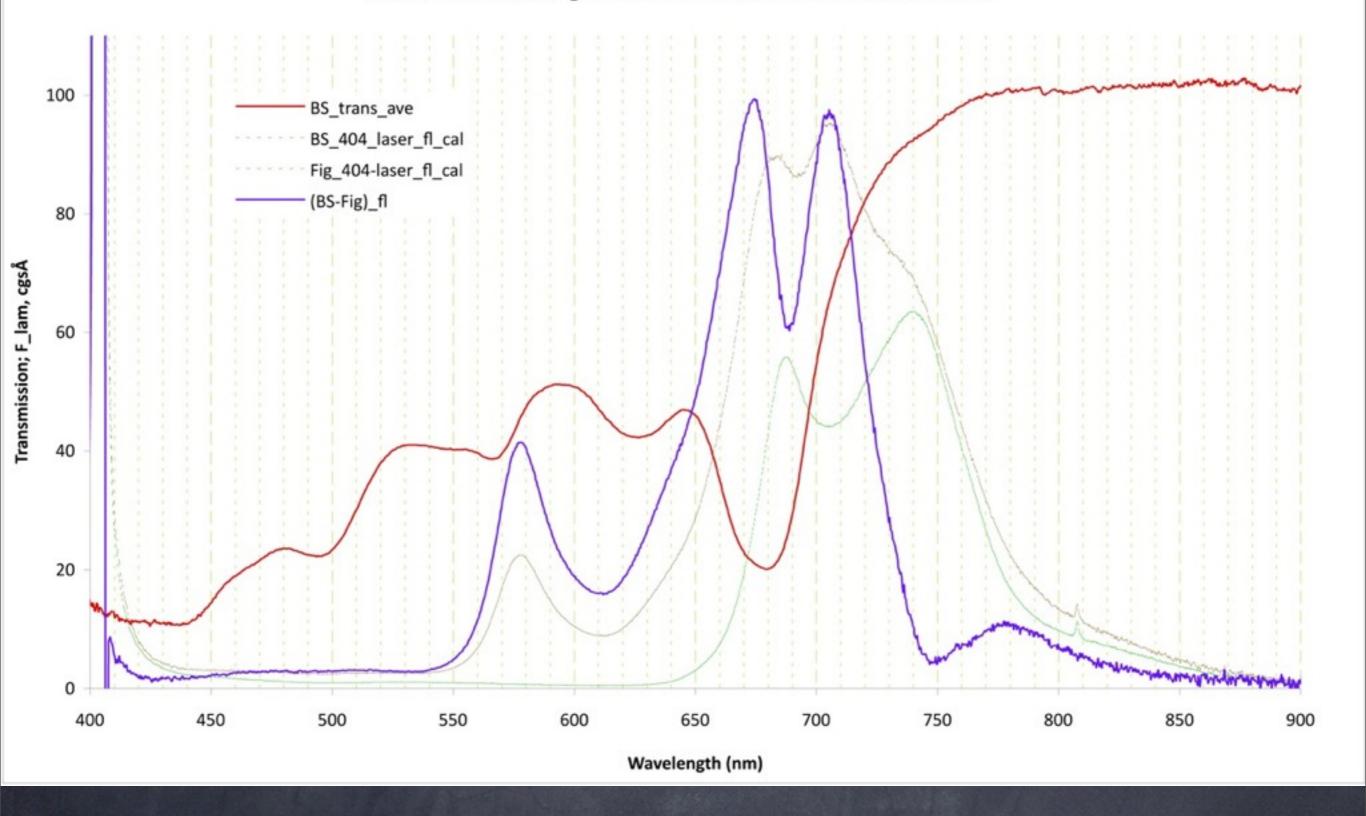
Hens egg, shell transmission and fluorescence, 404-laser



Chlorophyll, like haemaglobin, is a 'porphyrin' and there are other examples of this class that produce colours we see

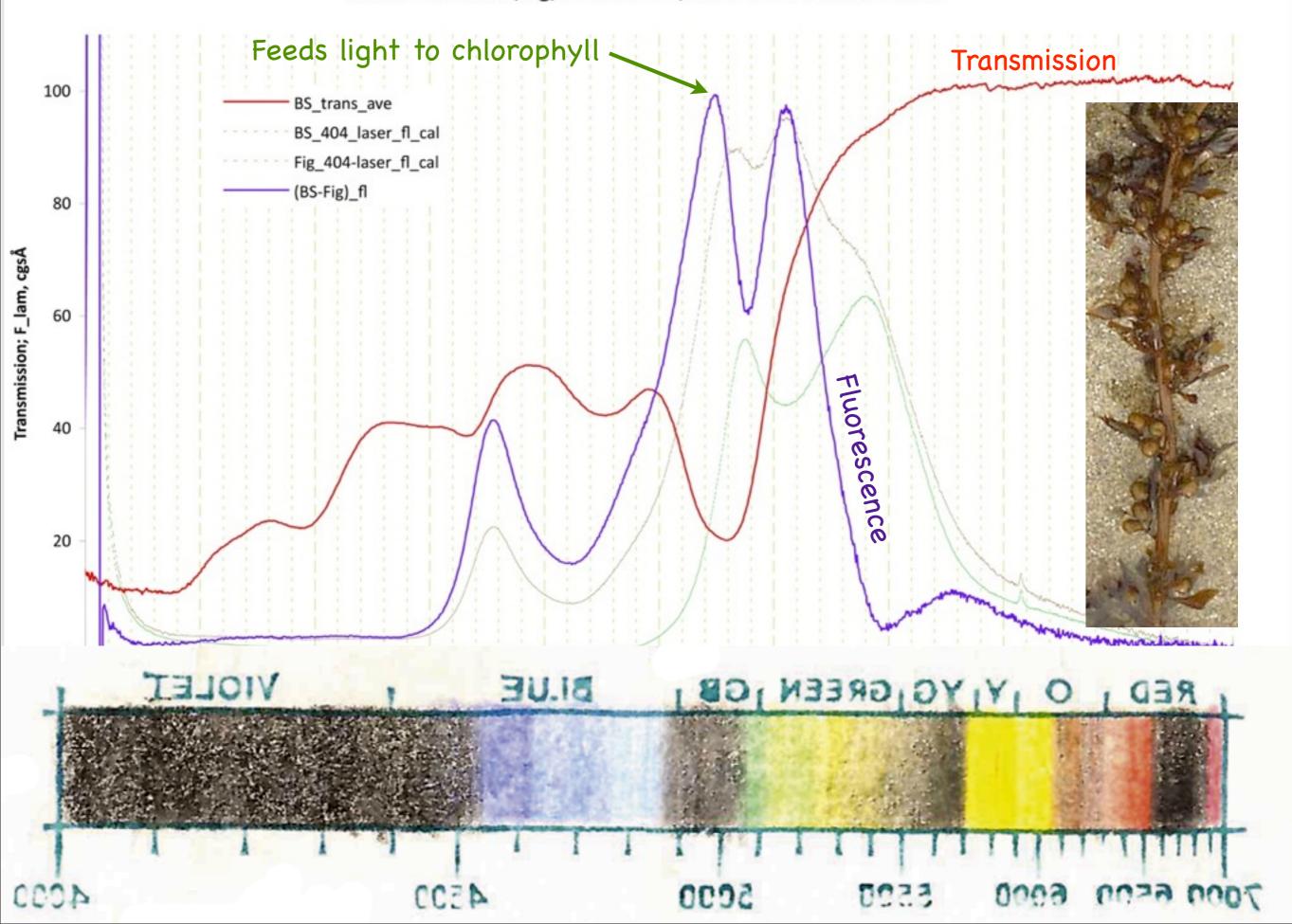
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Brown seaweed, Fig, fl and trans; RAEF 11 December 2010



Brown seaweed with phycoerythrin and chlorophyll

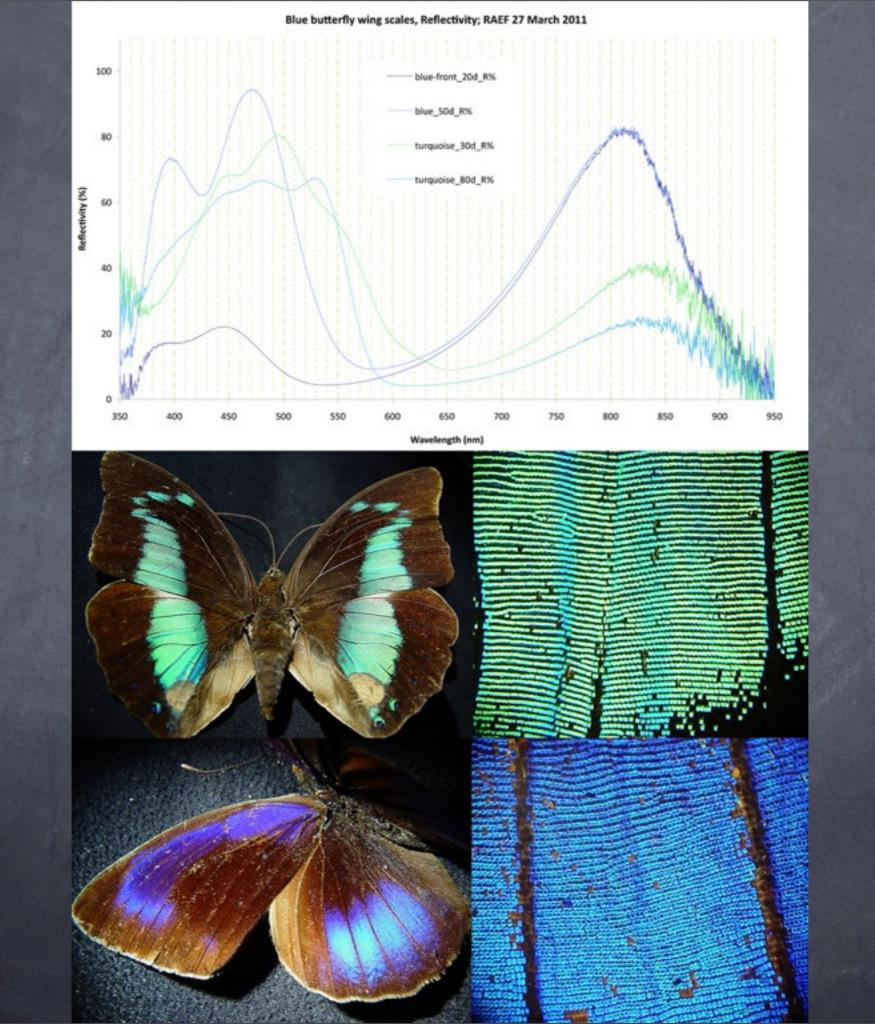
Brown seaweed, Fig, fl and trans; RAEF 11 December 2010

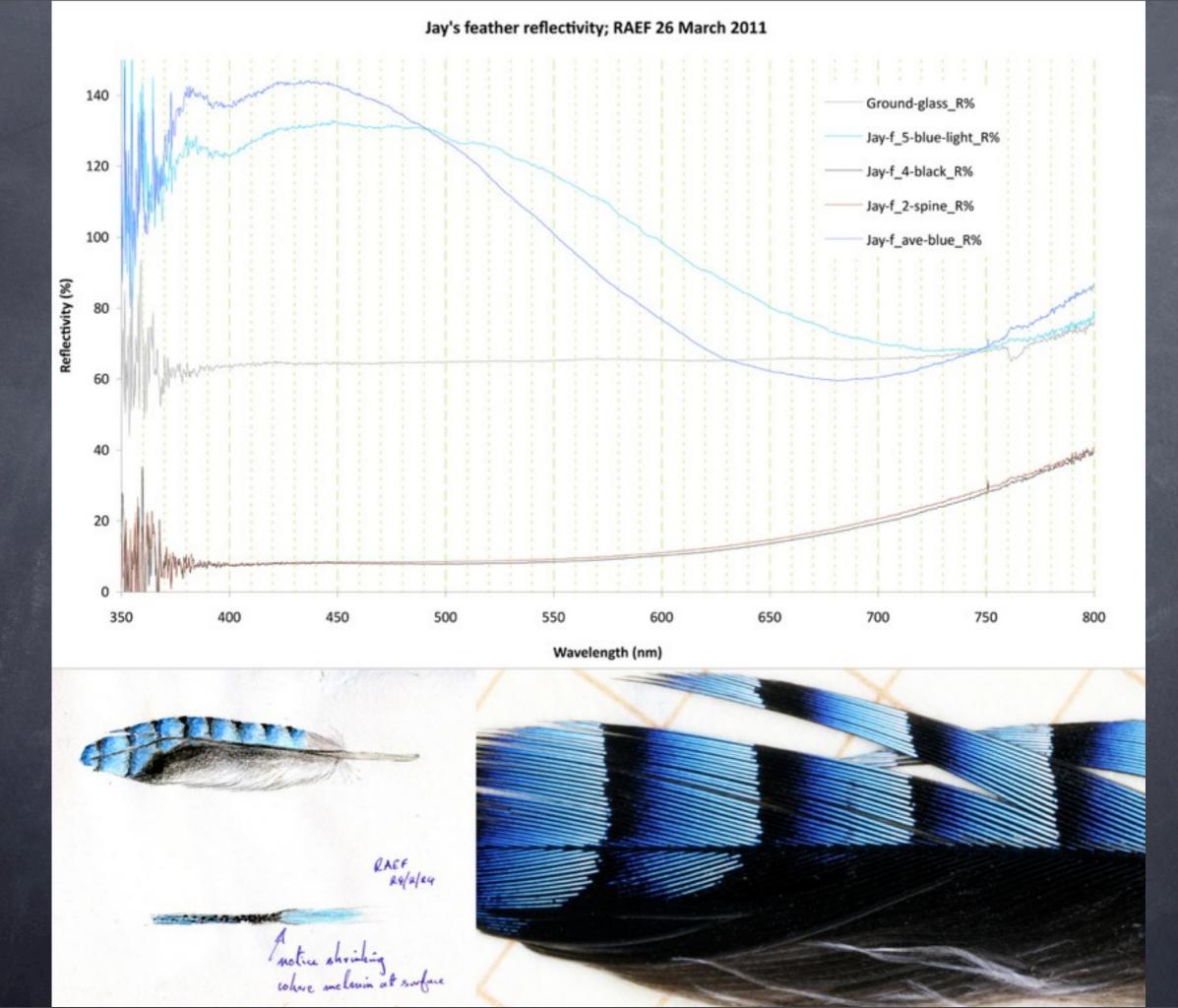


Animal blues

Many animal blues are not caused by pigments but are 'structural' — basically Tyndall (Rayleigh) scattering and/or interference phenomena

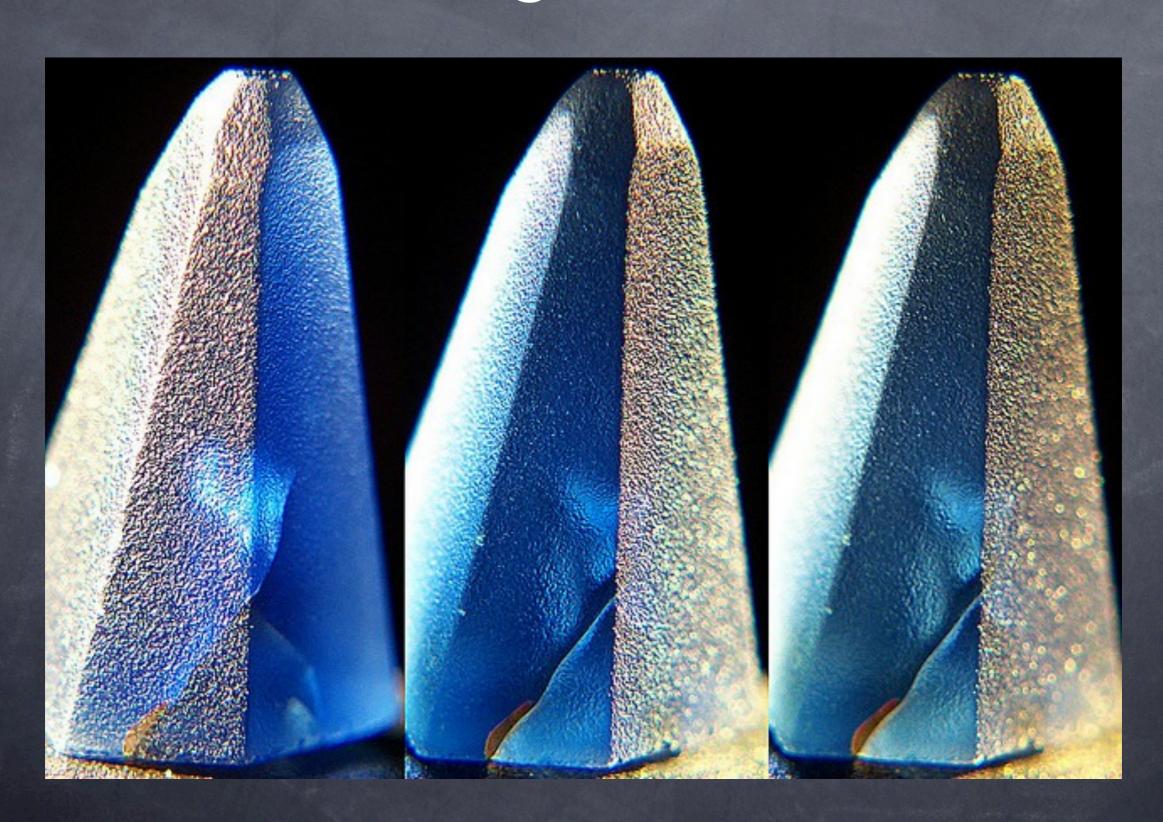
Animal blues





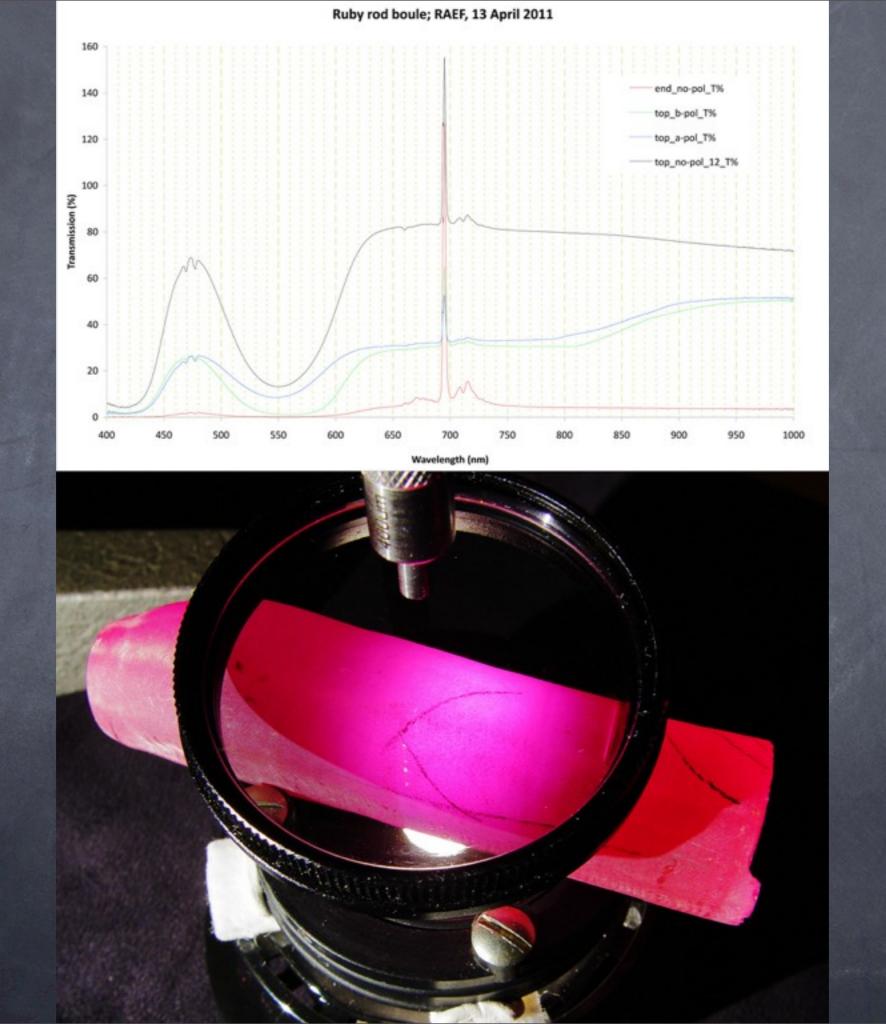
Richard Bingham. May's fouther found 1. light sky-blue 2. medium blue 3. derk blue 4. ter-black. light blue Uniform Tyndall scattlering surface of black melanin bloks.

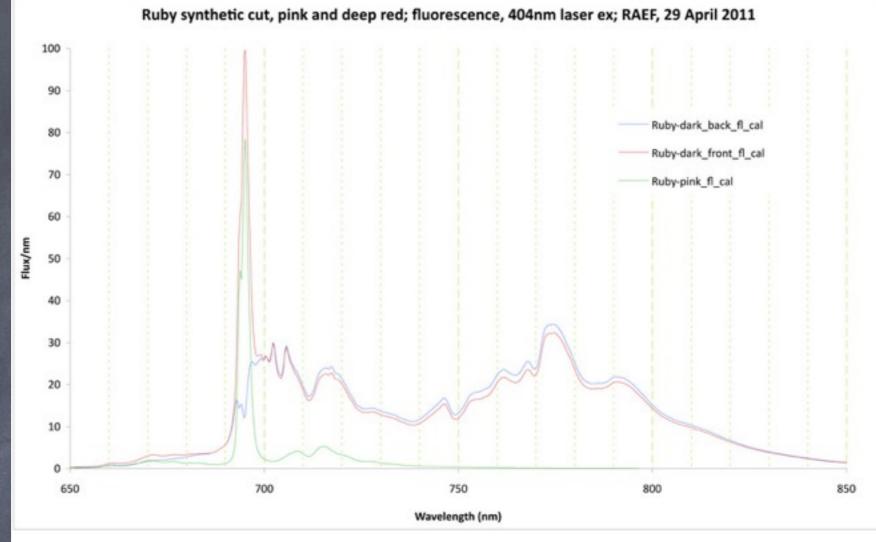
Solids (gemstones)



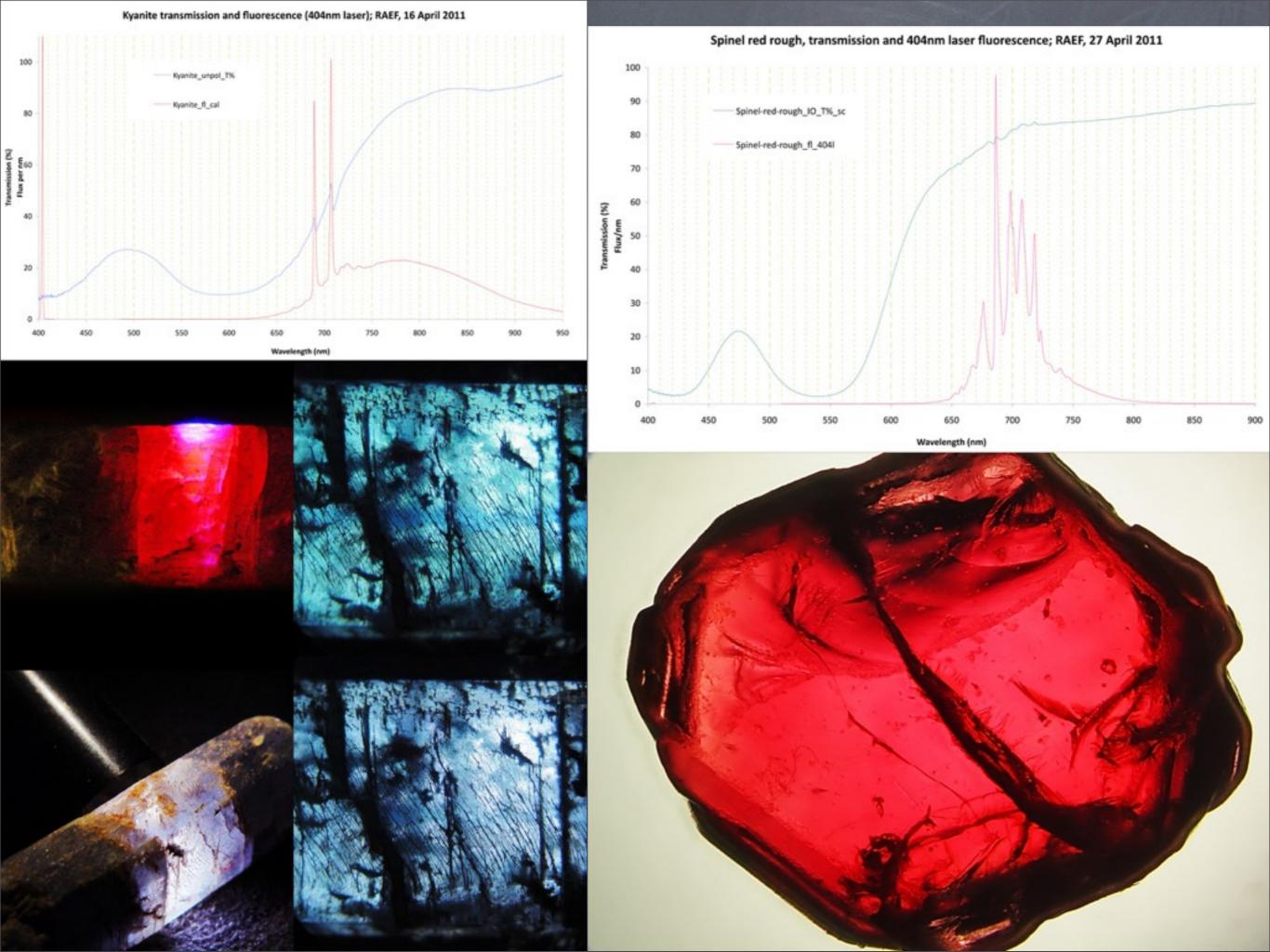
Solids (gemstones)

Different mechanisms produce colour in solids sometimes this is intrinsic to the basic material but is often due to 'impurities'. The most famous is the role of impurity chromium in ruby, emerald, spinel, kyanite etc. where the surrounding crystal structure generates a ligand field that allows optical transitions. Transitions from metastable levels can lase (the ruby laser)





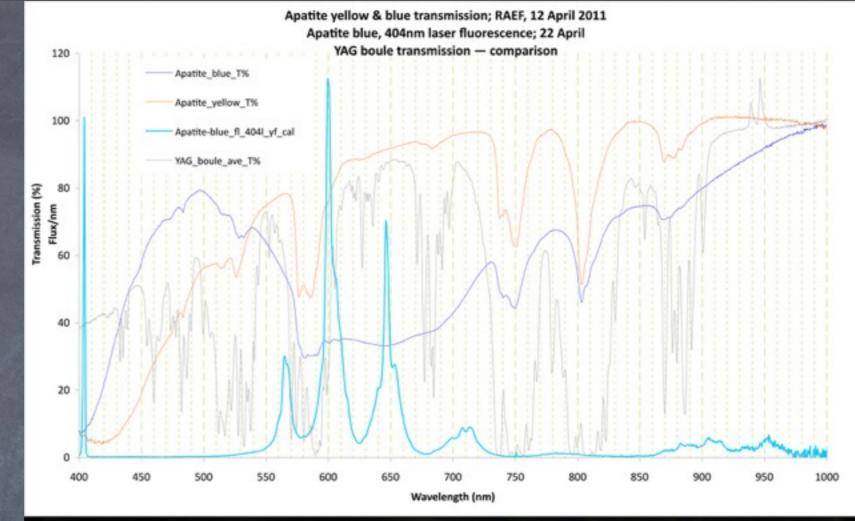




Apatite (calcium phosphate – Chloride Fluoride Hydroxide) contains rare earths (eg. neodymium) that make inner-shell transitions and produce rich absorption spectra. The Nd^3+ emits the strong NIR fluorescence lines that are exploited in Nd YAG lasers

The transmission spectrum of YAG is shown as the faint grey line

There is lots of apatite in your bones and teeth

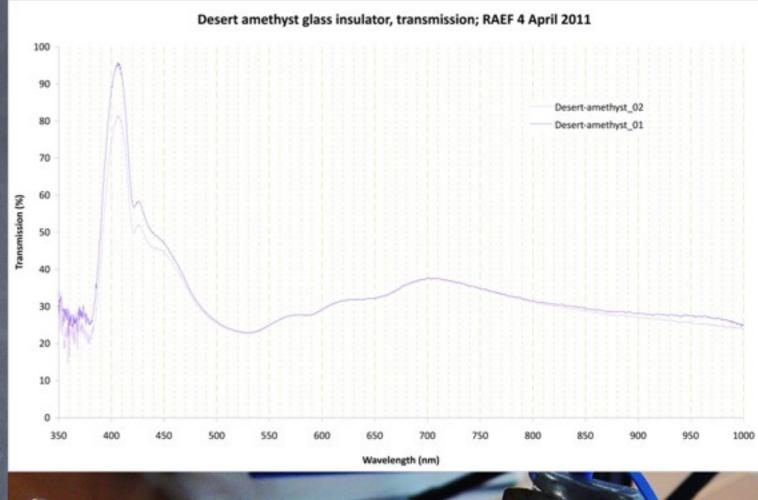




The effect of decades of desert sunlight on glass containing manganese and iron

This old glass telephone insulator was found on a fallen (unused!) telegraph pole by Trish on a trip from Parkes to Sydney

The colour (purple with high transparency in the blue) is from UV-induced colourcentres similar to those in adaptive spectacles







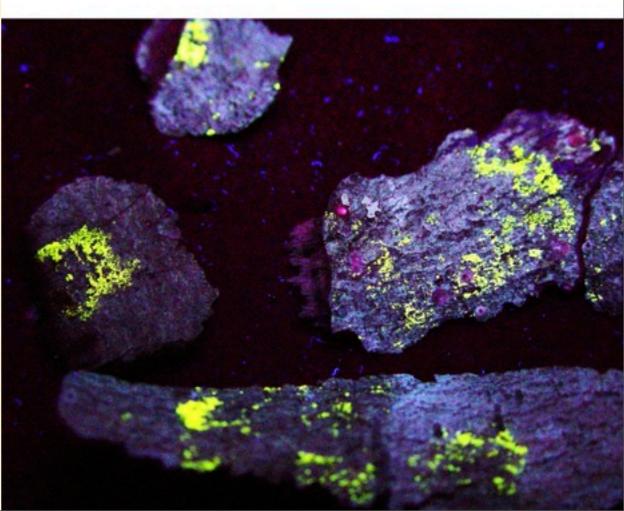












Trans. Dut The T. 1860, W. Wills.

The colours discussed in the first part of the talk are (and will be) important for characterising exo-Earths. It may be a while, however, before the later examples become of interest to astrophysicists (except to me of course!)

